

# Optical transitions in semiconductors

Semiconductor devices II - EE-567

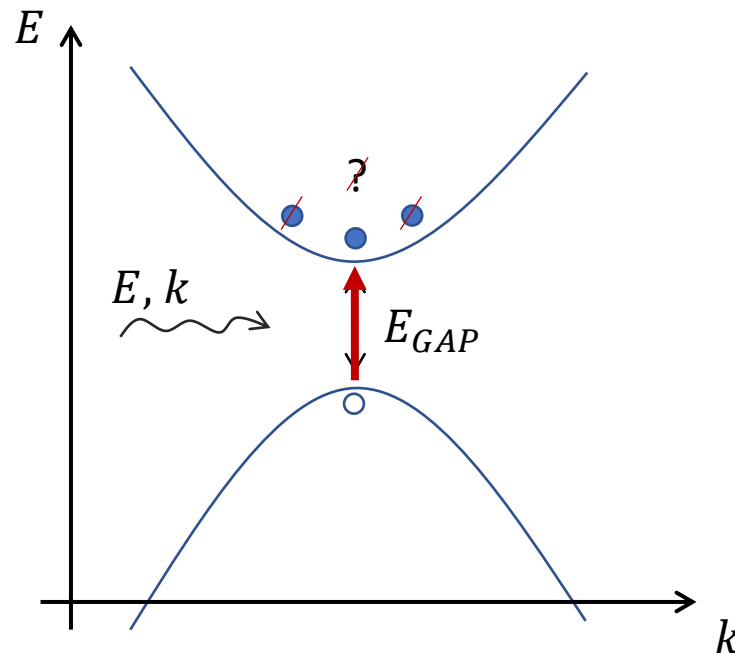
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# K-Selection Rule

- Conservation of momentum is true for semiconductors:  $\Delta p = 0$ ;
- Another way to write it :  $\Delta k = 0$  so  $k_0 = k_{final}$ ;

What this means for **direct bandgap semiconductor**



For optoelectronic devices we are interested in interactions between electrons and photons. Main points:

1. Energetic requirement for radiative interactions:  $E > E_{GAP}$
2. Conservation of momentum:

$$k_C = k_V + k_{ph}$$

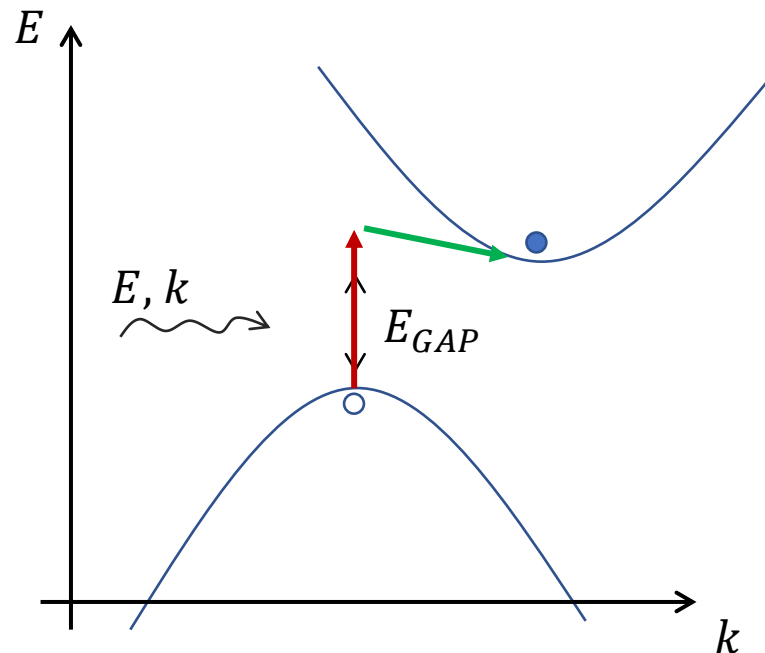
$$k_C, k_V = \frac{2\pi}{a} \approx 10^9 - 10^{10} cm^{-1}$$

$$k_{ph} = \frac{2\pi}{\lambda} \approx 10^4 cm^{-1}$$

# K-Selection Rule

- Conservation of momentum is true for semiconductors:  $\Delta p = 0$ ;
- Another way to write it :  $\Delta k = 0$  in total;

What this means for **in-direct bandgap semiconductor**



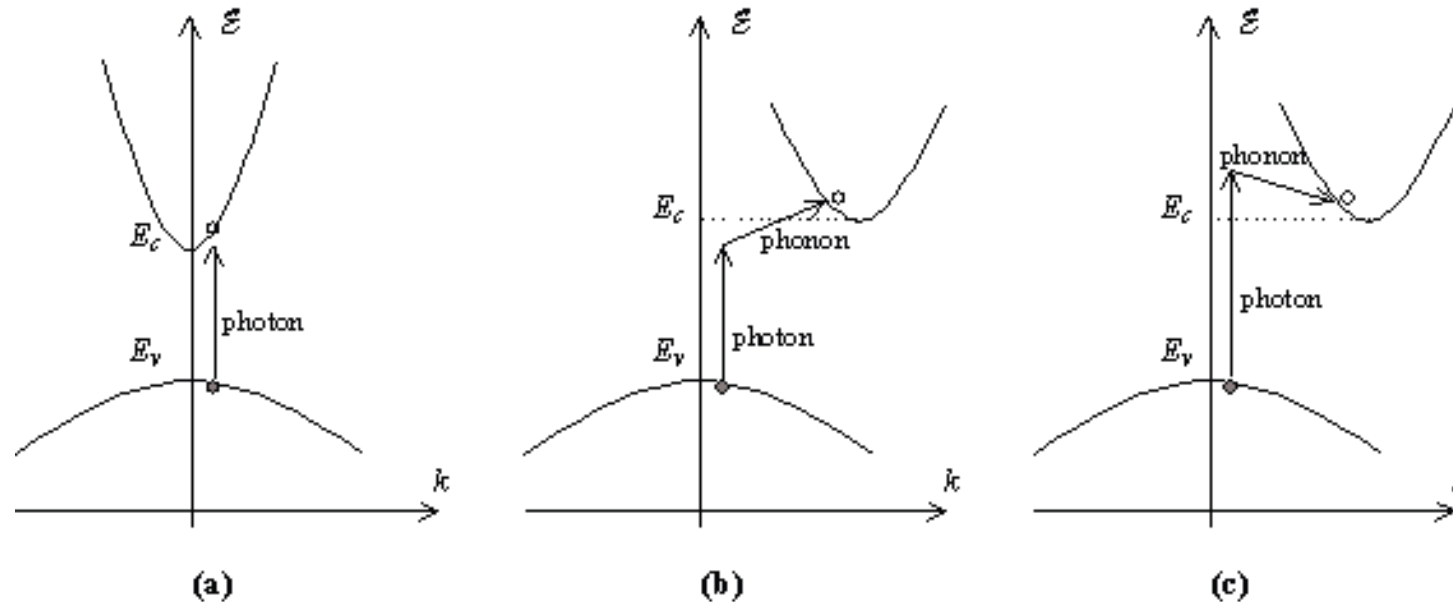
For optoelectronic devices we are interested in interactions between electrons and photons. Main points:

1. Energetic requirement for radiative interactions:  $E > E_{GAP}$
2. Conservation of momentum:

$$k_C = k_V + \cancel{k_{ph}} + k_{phonon}$$

Phonons have momentum comparable to that of electrons in the lattice

# Summary of Optical properties of semiconductors



How to quantify these microscopic phenomena?

## Direct Semiconductor

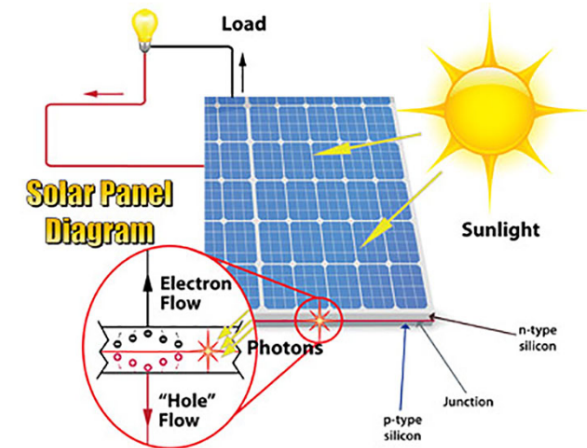
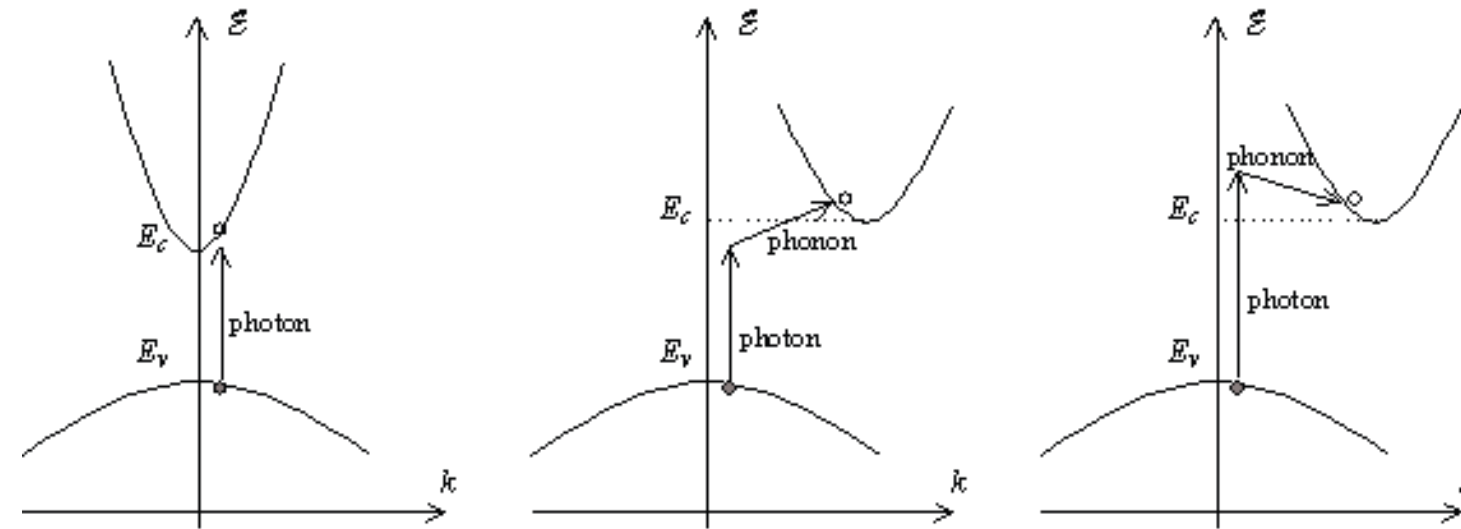
(a) Photon absorption in a direct bandgap semiconductor

## Indirect Semiconductor

(b) Photon absorption in an indirect bandgap semiconductor assisted by phonon absorption and

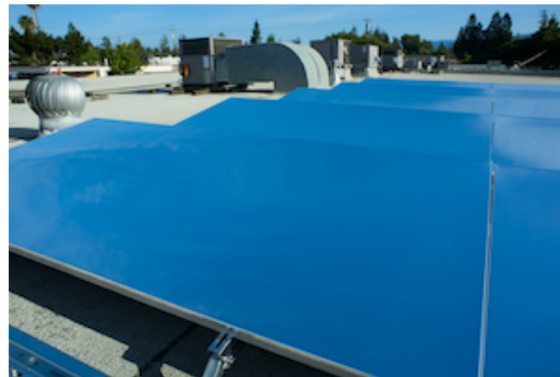
(c) Photon absorption in an indirect bandgap semiconductor assisted by phonon emission.

# Applications of Optical properties of semiconductors

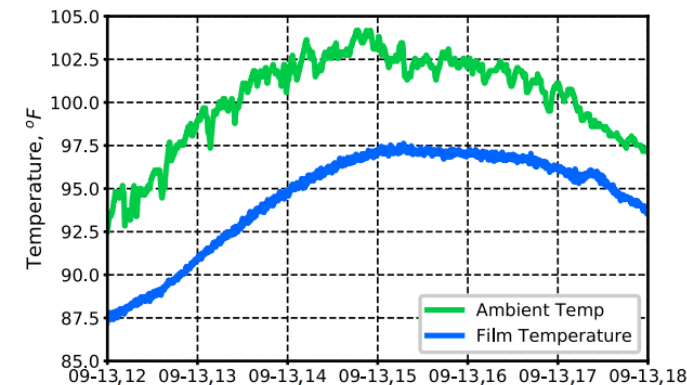


CASE STUDY  
Radiative Cooling Technology

**Harnessing the cold of the sky and space to enable electricity-free cooling**



(c)



Data measured from rooftop in Mountain View, CA on 9.13.2019

# Optical properties of semiconductors

$$H_0 = \frac{\mathbf{p}^2}{2m} + V(\mathbf{r})$$

In presence of an EM field described by a vector potential  $\mathbf{A}(\mathbf{r}, t)$  :

$$H = \frac{1}{2m} \left[ \mathbf{p} + \frac{e}{c} \mathbf{A}(\mathbf{r}, t) \right]^2 + V(\mathbf{r}) = H_0 + \frac{e}{mc} \mathbf{A} \cdot \mathbf{p} + \frac{e^2}{2mc^2} \mathbf{A}^2$$

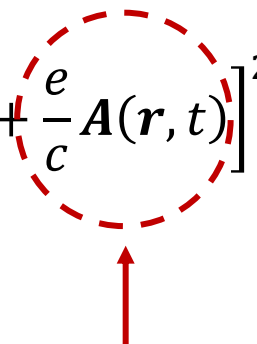
# Optical properties of semiconductors

$$H_0 = \frac{\mathbf{p}^2}{2m} + V(\mathbf{r})$$

Perturbation Theory



In presence of an EM field described by a vector potential  $\mathbf{A}(\mathbf{r}, t)$  :

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From the Lagrangian of a charged particle in a EM field, derived directly from Maxwells' equations


$$L = \frac{1}{2} m \mathbf{v} \cdot \mathbf{v} + eV + \frac{e}{c} \mathbf{v} \cdot \mathbf{A} \quad \longrightarrow \quad p_{tot} = \frac{\partial L}{\partial \mathbf{v}} = \mathbf{p} + \frac{e}{c} \mathbf{A}$$

# Optical properties of semiconductors

$$H_0 = \frac{\mathbf{p}^2}{2m} + V(\mathbf{r})$$

In presence of an EM field described by a vector potential  $\mathbf{A}(\mathbf{r}, t)$  :

The field is not extremely strong and  $e^2$  is small, thus this term can be neglected

$$H = \frac{1}{2m} \left[ \mathbf{p} + \frac{e}{c} \mathbf{A}(\mathbf{r}, t) \right]^2 + V(\mathbf{r}) = H_0 + \frac{e}{mc} \mathbf{A} \cdot \mathbf{p} + \frac{e^2}{2mc^2} \mathbf{A}^2$$


For example  $\mathbf{A}(\mathbf{r}, t) = A_0 \hat{\mathbf{e}} e^{i(\mathbf{q} \cdot \mathbf{r} - \omega t)} + c.c.$  with  $\hat{\mathbf{e}} \perp \mathbf{q}$

$$H = H_0 + \frac{eA_0}{mc} e^{i(\mathbf{q} \cdot \mathbf{r} - \omega t)} \hat{\mathbf{e}} \cdot \mathbf{p} + \frac{eA_0}{mc} e^{-i(\mathbf{q} \cdot \mathbf{r} - \omega t)} \hat{\mathbf{e}} \cdot \mathbf{p}$$

# Optical properties of semiconductors

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Absorbed photon

Emitted photon

# Fermi golden rule

Coupling between initial and final states

Density of states in the transition

$$W_{ij} = \frac{2\pi}{\hbar} |\langle i | H' | j \rangle|^2 \rho(E_f)$$

$W_{ij}$  is the *transition probability per unit time* between two quantum states  $i, j$

$\rho(E_f)$  is the density of final states. If energy is conserved:  $\rho(E_f) = \delta(E)$

$H'$  is the perturbation term in the Hamiltonian

$|\langle i | H' | j \rangle|$  is the matrix element for the interaction

In our case:

$$W_{i \rightarrow j} = \frac{2\pi}{\hbar} \left( \frac{eA_0}{mc} \right)^2 |\langle \psi_j | e^{i\mathbf{q} \cdot \mathbf{r}} \hat{\mathbf{e}} \cdot \mathbf{p} | \psi_i \rangle|^2 \delta(E_j - E_i - \hbar\omega)$$

$$W_{j \rightarrow i} = \frac{2\pi}{\hbar} \left( \frac{eA_0}{mc} \right)^2 |\langle \psi_i | e^{-i\mathbf{q} \cdot \mathbf{r}} \hat{\mathbf{e}} \cdot \mathbf{p} | \psi_j \rangle|^2 \delta(E_i - E_j + \hbar\omega)$$

Absorbed photon  
 $E_j = E_i + \hbar\omega$

Emitted photon  
 $E_i = E_j - \hbar\omega$

# Transitions

We have many  $i, j$  states from and to which the transition can happen

The net transition rate from  $i$  to  $j$  is given by  $W_{ij} - W_{ji}$ . Over all states:

$$W = \sum_{ij} W_{ij} - W_{ji}$$

← Considering photons interacting with a semiconductor, it represents the net absorption rate

Every state has an occupation probability given by  $f(E)$

To have a transition between  $i$  and  $j$ , we need occupied  $i$  and empty  $j$ :  $f(E_i)[1 - f(E_j)]$

So, we multiply  $W_{ij}$  by  $f(E_i)[1 - f(E_j)]$  and  $W_{ji}$  by  $f(E_j)[1 - f(E_i)]$ :

$$W = \frac{2\pi}{\hbar} \left( \frac{eA_0}{mc} \right)^2 2 \sum_{ij} |\langle \psi_j | e^{i\mathbf{q} \cdot \mathbf{r}} \hat{\mathbf{e}} \cdot \mathbf{p} | \psi_i \rangle|^2 \delta(E_j - E_i - \hbar\omega) [f(E_i) - f(E_j)]$$

# Transitions

$$W = \frac{2\pi}{\hbar} \left( \frac{eA_0}{mc} \right)^2 2 \sum_{ij} |\langle \psi_j | e^{i\mathbf{q} \cdot \mathbf{r}} \hat{\mathbf{e}} \cdot \mathbf{p} | \psi_i \rangle|^2 \delta(E_j - E_i - \hbar\omega) \underbrace{[f(E_i) - f(E_j)]}_{= 1 \text{ at } T = 0}$$

In the limit for  $T = 0$ , full VB and empty CB:

$$W \sim \sum_{\text{occ. } i} \sum_{\text{free } j} |\langle \psi_j | e^{i\mathbf{q} \cdot \mathbf{r}} \hat{\mathbf{e}} \cdot \mathbf{p} | \psi_i \rangle|^2 \delta(E_j - E_i - \hbar\omega)$$

$$\sim \sum_{cv} \int_{BZ} |\hat{\mathbf{e}} \cdot \langle \psi_{c\mathbf{k}} | \mathbf{p} | \psi_{v\mathbf{k}} \rangle|^2 \delta(E_{c\mathbf{k}} - E_{v\mathbf{k}} - \hbar\omega) \frac{d\mathbf{k}}{(2\pi)^3}$$

**Perturbation theory**  $\rightarrow$  we can apply our knowledge on the wavefunctions in the case without perturbations and solve the calculation of the net absorption rate

# Joint density of states

The dipole matrix elements are smooth functions of the  $\mathbf{k}$ -vector.  
Thus, they do not affect much the integral, and can be factorized out.

On the contrary, the main term to determine the behaviour of the calculation is  $\delta(E_{c\mathbf{k}} - E_{v\mathbf{k}} - \hbar\omega)$

Essentially, the integral is controlled by the so-called *joint density of states*:

$$\rho_{cv}(\omega) = \int \delta(E_{c\mathbf{k}} - E_{v\mathbf{k}} - \hbar\omega) \frac{d\mathbf{k}}{(2\pi)^3} \quad \rightarrow \quad \rho_{cv}(\omega)^{3D} = 2\pi \left( \frac{2m^*}{\hbar^2} \right)^{\frac{3}{2}} \sqrt{\hbar\omega - E_g}$$

Specific case of a **direct** bandgap semiconductor in **3D**

# Dielectric function and transition rate

The *total energy per unit time* dissipated in the system (volume  $V$ ) is:

$$P(q, \omega) = \hbar\omega W(q, \omega)$$

From Maxwell equations we know that

$$\left. \begin{aligned} \frac{\partial D}{\partial t} &= \frac{\partial E}{\partial t} + 4\pi J \\ D &= \epsilon_1 + i\epsilon_2 \\ J &= \sigma \cdot E \end{aligned} \right\} \rightarrow \epsilon = 1 + \frac{4\pi i\sigma}{\omega}$$

$$\hbar\omega W = \frac{\omega}{4\pi i} (\epsilon - 1) |E_0|^2 V$$

Also, from general electromagnetism:

$$P = \int_V J \cdot E \, dr = \int_V \sigma E \cdot E \, dr = \sigma |E_0|^2 V$$

Thus we connect the microscopic description to the macroscopic quantities  $\sigma, \epsilon$

# Absorption coefficient

In the end we find:

$$\epsilon_2(\mathbf{q}, \omega) = \frac{2\pi\hbar c^2}{\omega^2} \frac{1}{V} \frac{W(\mathbf{q}, \omega)}{A_0^2}$$

$$\alpha(\omega) = \frac{2\pi\hbar c}{n(\omega)\omega} \frac{1}{V} \frac{W(\omega)}{A_0^2}$$

$$I(z) = I_0 e^{-\alpha(\omega)z}$$

Familiar **macroscopic** absorption law

# Absorption coefficient

Direct Bandgap in 3D:

$$\alpha \approx \sqrt{h\nu - E_G}$$

Indirect Bandgap:

$$\alpha \propto \frac{(h\nu - E_G - E_P)^2}{e^{\frac{E_P}{kT}} - 1} - \frac{(h\nu - E_G - E_P)^2}{1 - e^{-\frac{E_P}{kT}}}$$

Direct Bandgap is more efficient for photon absorption: MoS<sub>2</sub>, GaAs, InP...